# Astronomy 5465 Galaxies

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## Astro 5465 Tuesday. Jan. 18, 2022 Today's Topics

- Class Overview
  - Overview of Material
  - Syllabus
  - Schedule
- Introductions
  - Who are you?
  - Who am I?
- Course Prologue
  - Summary of Galactic and Extragalactic Astronomy
  - Historical Overview

## **Highlights of the Syllabus**

#### • Text:

- Galaxies in the Universe – Sparke & Galagher

#### Additional References:

- Galactic Astronomy Binney and Merrifield
- Galactic Dynamics Binney & Tremaine

#### • Popular Texts:

- Coming of Age in the Milky Way Ferris
- Lonely Hearts of the Cosmos Overbye
- Minding the Heavens: Story of Our Discovery of the Milky Way Belkora

#### Additional Readings:

- Articles & Papers from the Literature, especially Annual Reviews
- Lectures: Reading done in advance, notes on the web
  - One chapter covered each week!
- Homework: Typically Assigned/Due on Wednesdays
- Exams: Midterm Exam + Final Presentation
- Grading: Exams 50%, Homework 50%

## **Syllabus Continued**

- This Course will Feature Three Themes:
- Galaxies as stellar population factories
- Galaxies as fossil remnants of early formation processes
  - Secular evolution, evolution of stellar populations, etc.
- Sites of unique astrophysics (dark matter)
- Lecture:
  - I plan to lecture 3 hours per week but let each of you lead a discussion of the assigned readings (~ 15 min. on Friday).
- Homework (25%):
  - Homework will be assigned periodically: analytic & computational
- Exams (50%):
  - Midterm (25%) and Final Presentation (25%)

## Who am I?

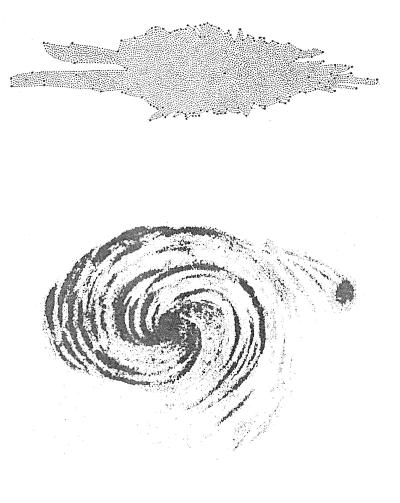
#### • Background

- Ph.D University of Hawaii
  - Measured Expansion of the Universe
  - Inferred Existence of Dark Energy
- Plaskett Fellow, Herzberg Institute for Astrophysics (Victoria, BC)
- Research Fellow, Kitt Peak National Obs. (Tucson, AZ)
- Assistant Professor, Indiana University
- Associate Professor, University of Wyoming
- Research Interests
  - Evolution of Galaxies
  - Gravitational Lensing & Cosmology
  - Astronomical Instrumentation

### What Do You Think of When You Think of Galaxies?

### **Brief History of Galaxies**

- Galileo Uniform distribution of Bright Stars but Faint Stars form Flattened Plane -Milky Way
- Kant Galaxies as Island Universes
  - Solar System analogy: flattened structure dominated by gravitational force
  - The Nebulae could be similar but very distant systems
- Messier's Catalog of the brightest nebulae (star clusters & galaxies)
- Herschel's catalogs of nebulae (circa 1780) followed by Dreyer's NGC (New General Catalog)
- Herschel's star counts (#stars vs. mag.) reveal flattened disk of stars
- Ross (circa 1845) observes spiral structure in some nebulae

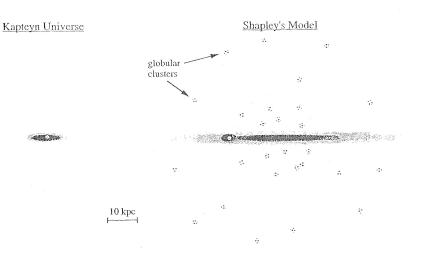


### **Brief History of Galaxies Cont.**

- Kapteyn Selected Areas around the sky (~200 designated by IAU)
  - World-wide effort to count stars vs. mag. (pg mags: calibration wrt North Polar Sequence)
  - Spectral Types and velocities of brightest stars
  - Results (collected by Kapteyn [1909, Ap.J. 29, 46, 30, 284, 398]):
    - MW is a flattened distribution of stars (5:1 axial ratio)
    - Sun at center with density decrease in all directions (50% at 800 pc, 10% at 2800 pc)
    - Sun about 650 pc from center
    - Kapteyn considered extinction but since reddening was modest concluded it was negligible (assumed scattering dominated, not extinction)
- Trumpler (1930) discovers significant extinction using Galactic clusters
  - stars in clusters dim faster than distance (angular size) would produce
  - star counts in "dark nebulae" are offset wrt unobscurred regions

### **Brief History of Galaxies Cont.**

- Shapley (1918, 1919) discovers the true size of the Milky Way
  - distribution of Globular Clusters not centered on Sun
    - centered about 15 kpc from Sun (RR Lyrae variables)
    - concentrated in direction of Sagittarius
    - Milky Way is 100 kpc across
    - Recognized "zone of avoidance" but didn't connect it with extinction
- Shapley-Curtis Debate (1920) on
  - Size of the Milky Way and Location of the Sun
  - Distance and Nature of Spiral Nebulae
- Shapley Galaxy is very big and contains spiral nebulae
- Curtis Galaxy is small and spiral nebula are distant galaxies
- Oort Suggested Extinction



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### **Brief History of Galaxies Cont.**

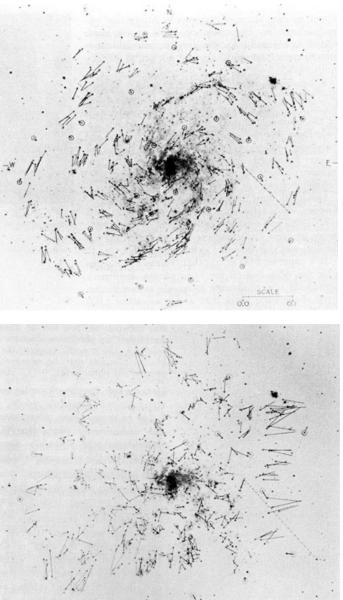
- Summary of the Great Debate on the Nature of Spiral Nebulae
- Both were right and both were wrong! Details:

•	<u>Curtis:</u>	<u>Shapley:</u>
•	Kapteyn: right (MW ~ 10 kpc)	wrong (MW ~ 100 kpc)
•	SN1885: Nova – M31 at 150 kpc	Nova – M31 at 150 kpc
•	(external galaxy)	(M31 inside MW!)
•	van Maanen's rotation of M33: ??	if V ~ 200 km/s, D < 50 kpc
•	high v stars: unbound to MW, why?	OK since MW is massive
•	MW appearance: like edge-on nebulae	??
•	Zone of Avoidance: neglected extinction	destruction of globulars
•	Hubble's Measurements of Extragalactic Cepheids	
	Hubble resolves M31 disk into stars (1920)	
	<b>Cepheid Period-Luminosity Relation (Leavitt 1920)</b>	

- Finds Cepheids in NGC 6822 and M31 (1922) and shows they are ~ 300 kpc away

#### Van Maanen vs. Lunmark

- Van Maanen claimed to have measured proper motion of rotation within M33
  - Most talented and respected astrometrist of his day
  - Work was unquestioned since it demonstrated what most expected.
- Lundmark checked this after Hubble's measurement of extragalactic Cepheids
  - Unable to confirm van Maanen's results
  - No evidence for rotation.
  - Huge embarrassment for Van Mannen
- This is an excellent example of the scientific method.
  - Best theoretical expectation of the day
  - Initially confirmed
  - New data led to inconsistency
  - Ultimately overturned.



#### **Early Ideas Regarding Stellar Populations**

- Lindblad (1927) stars move under mutual attraction of gravity
  - MW is an axisymmetric disk with 2 components:
  - One component with rotational motion
  - One component with random motion (explained high velocity stars)
- Oort (1927,28) high velocity stars form a distinct kinematic component, but most in circular orbits traveling with the Sun
- Baade (1944, Ap.J. 100, 137, 147) develops concept of two distinct stellar populations
  - Resolves bulge of M31 into stars (i.e., Tip of Red Giant Branch)
  - Spiral Galaxies Also Composed of Two Populations (just like MW)
- Pop. I Blue and Red Supergiants
  - Disk-like structure with rotational motion
  - Young and metal rich (Solar comp.) but some older stars (Sun-like) too
- Pop. II Red Giants and dwarfs
  - Spherical-like structure with random motions
  - old, metal poor (no young metal-poor stars or old metal-rich stars)
  - MW bulge? No! not associated with the local metal-poor stars (the halo)
- Pop. III (extremely low metallicity but what about bulges)?

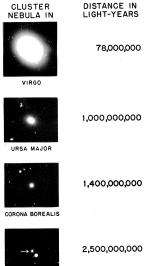
#### **Optical Classification of Other Galaxies**

#### • Hubble Morphological Sequence

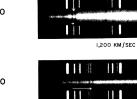
Nebulae are Galaxies but so different
E's -> S0s -> Sa -> Sb -> Sc with parallel barred sequence

- Structural and Color Trends
  - -Yes, red and blue sequences
  - -Populations I & II? (No!)
- Kinematical Difference
  - -Rotation vs. Velocity Dispersion
- Velocity-Distance Relation

-Universe expanding (more later)

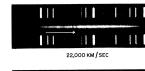


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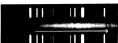
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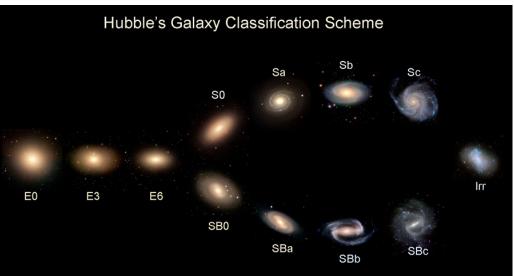
RED-SHIFTS

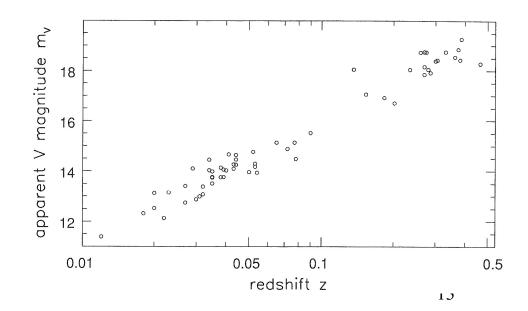




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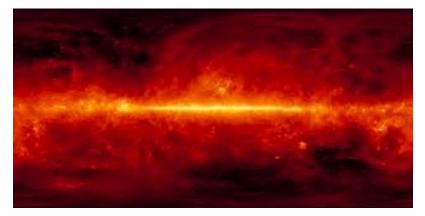


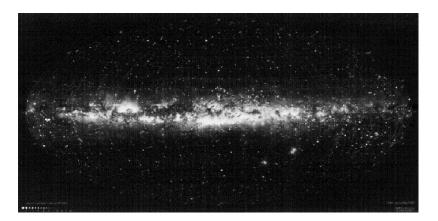


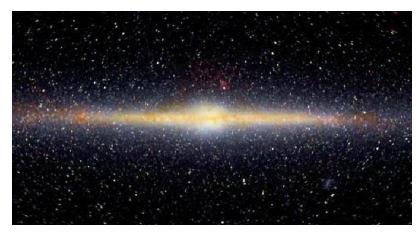
BOOTES

## Milky Way is a Disk Galaxy

- Optical Morphology Reveals Disk of Stars and Dust
  - HII Regions, Star Clusters, etc.
  - Bulge Seen Through "Baade's Window"
- Near Infrared Reveals Near Dust Free Distribution of Stars
  - Fall-off in Stellar Density with Radius
  - Full Bulge and it's "Boxy" Morphology
- Far Infrared Reveals Cool Dust



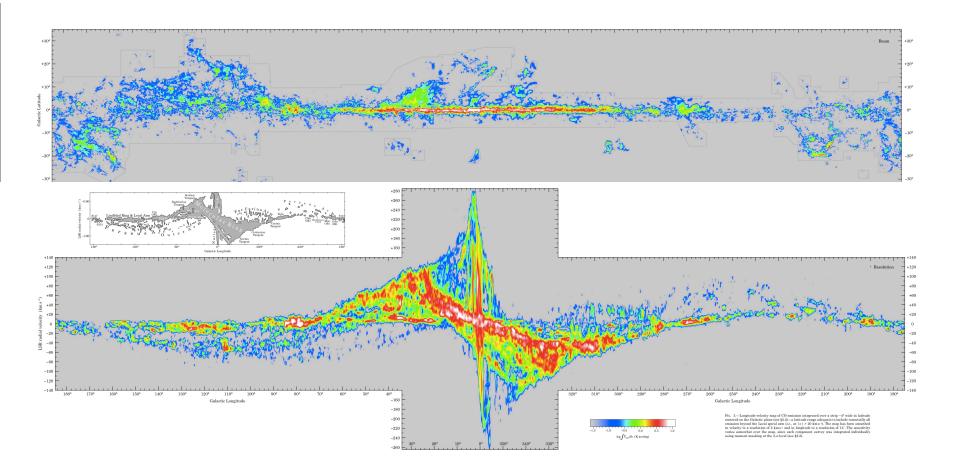




http://astrog80.astro.cf.ac.uk/Chromoscope-tiles/

#### **Interstellar Medium of the Milky Way**

- Jansky (1932) discovers radio continuum of the Milky Way
- Van de Hulst (1944) uses quantum mechanics to predict the 21-cm line (hyperfine) line of neutral Hydrogen
- Oort et al. (1958) map the MW at 21-cm (maximum velocity = tangent point, more later)
- Recent Maps of MW in CO (Dane et al. 2001)

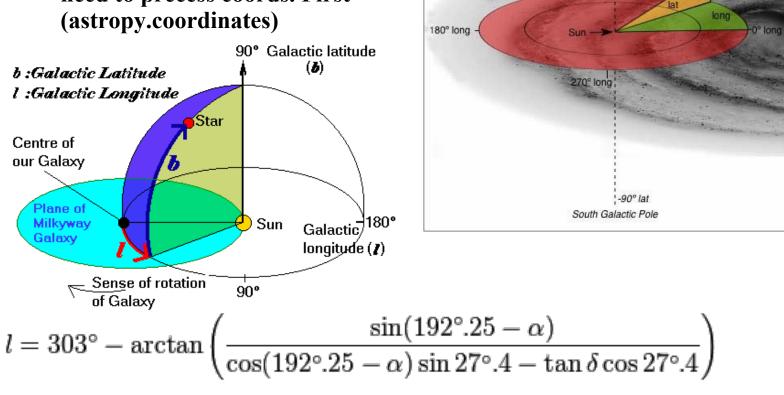


## **Galactic Coordinate System**

North Galactic Pole

90° long

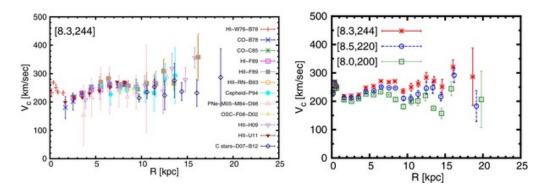
- Galactic Coordinates are Defined wrt The Center of the Galaxy
  - Transformations given below are in 1950 coords. So you need to precess coords. First (astropy.coordinates)

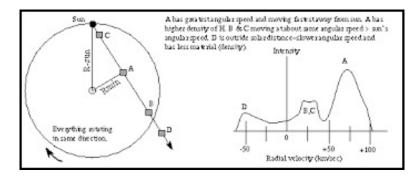


 $\sin b = \sin \delta \sin 27^{\circ}.4 + \cos \delta \cos 27^{\circ}.4 \cos(192^{\circ}.25 - \alpha)$ 

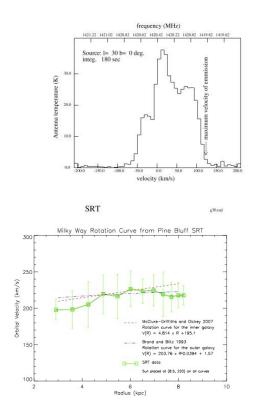
### **Milky Way HI Rotation Curve**

- MW HI Easily Detected as 21-cm Emission
- Along Any Line-of-Sight HI Has Range of Velocity
  - Clouds, Complex Structure
  - Maximum Velocity Found Near Tangent Point Due to Orbital Projection
  - Nice Data from Student Led Projects (see Figure Spectrum from Haystack Obs.)
  - Review: Bhattacharjee 2014 (ApJ 785, 63)









# **Reading this Week**

#### By Thursday:

**Start Reading Chapter 1 in text:** (review of stellar properties)

HW #1 (due Tues. Jan 25):

**a)** Use the data summarized by Bhattacharjee (2014) to plot the rotation curve in Python. Use your own judgement on how to plot it.

b) Compute an appropriate average by fitting a spline to each data set for interpolation and combining them.

c) Assume a spherical mass model for the Galaxy and compute both M(R) and  $\rho(R)$ . Make a plot of your results and fit a polynomial model. Comment on your results and provide a summary of your results.