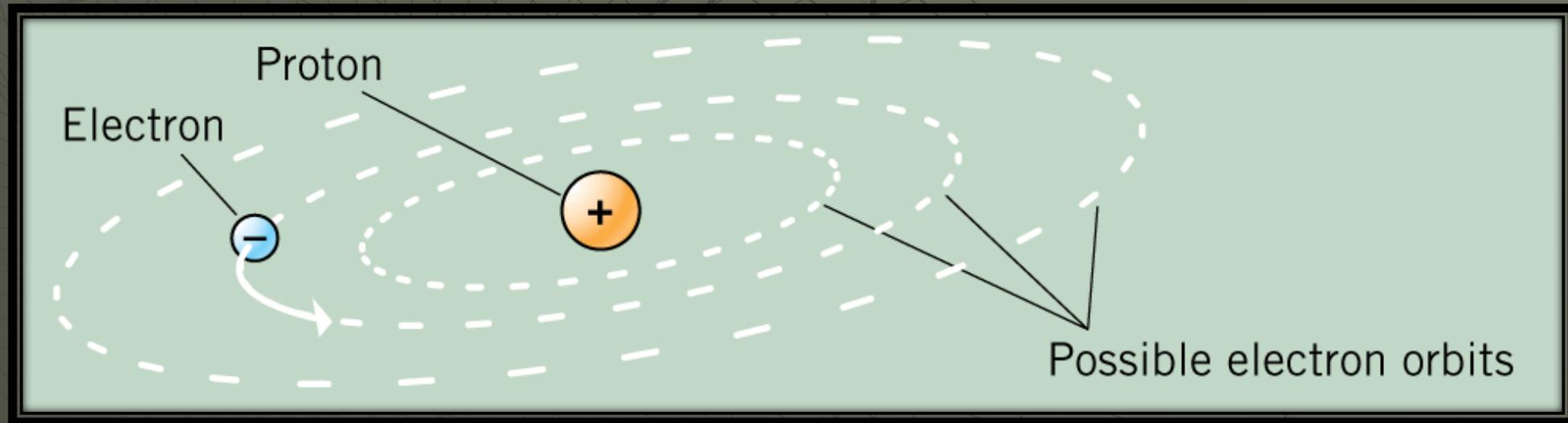


Atoms & Spectra

- ◆ Objectives
 - How does light interact with elements
 - Different types of spectra

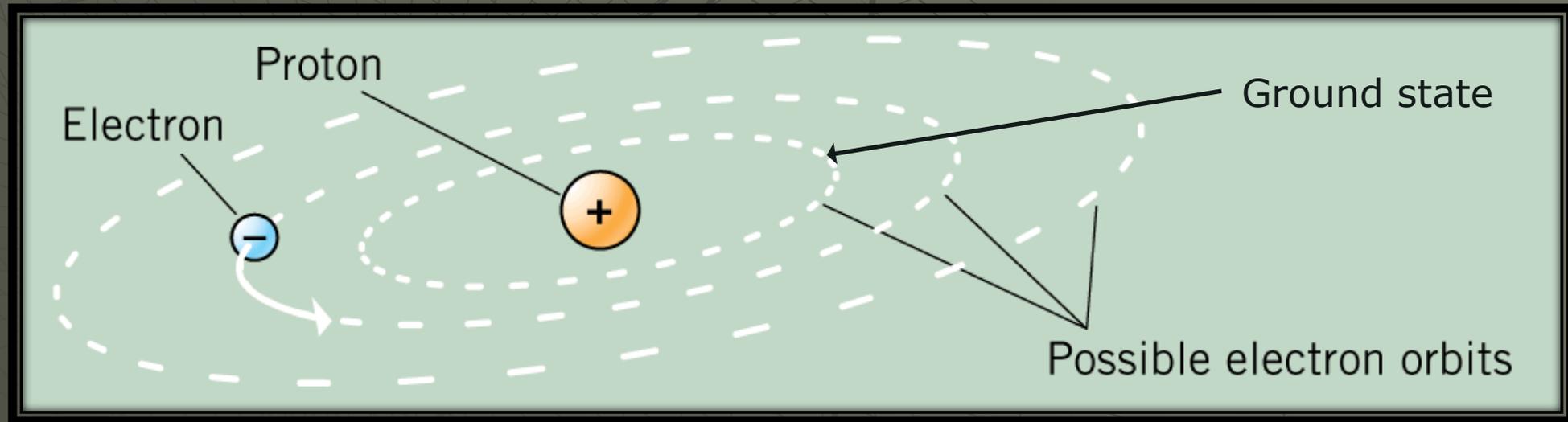
Bohr Model of Atom

- ◆ **Hydrogen** atom
 - 1 proton (nucleus), and 1 electron (orbit)
- ◆ Proton mass/electron mass $\sim 2000!$
- ◆ **Will light interact with the electron or the proton?**



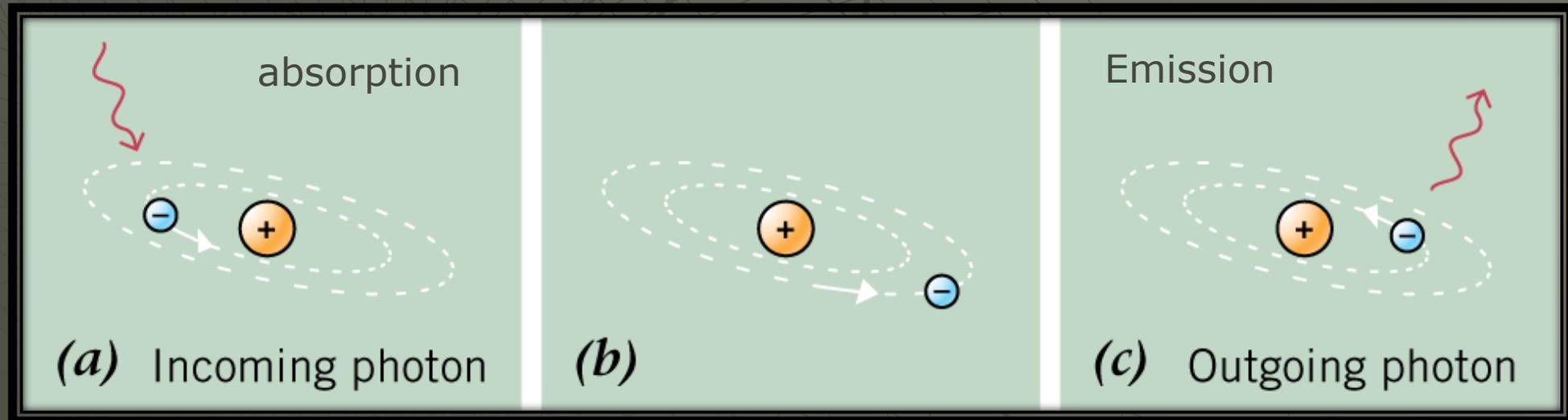
Bohr Model of Atom

- ◆ Electron “orbits” the proton
- ◆ Has very specific stable “orbits”
- ◆ Smallest orbit is called the **ground state**
 - **Lowest energy**



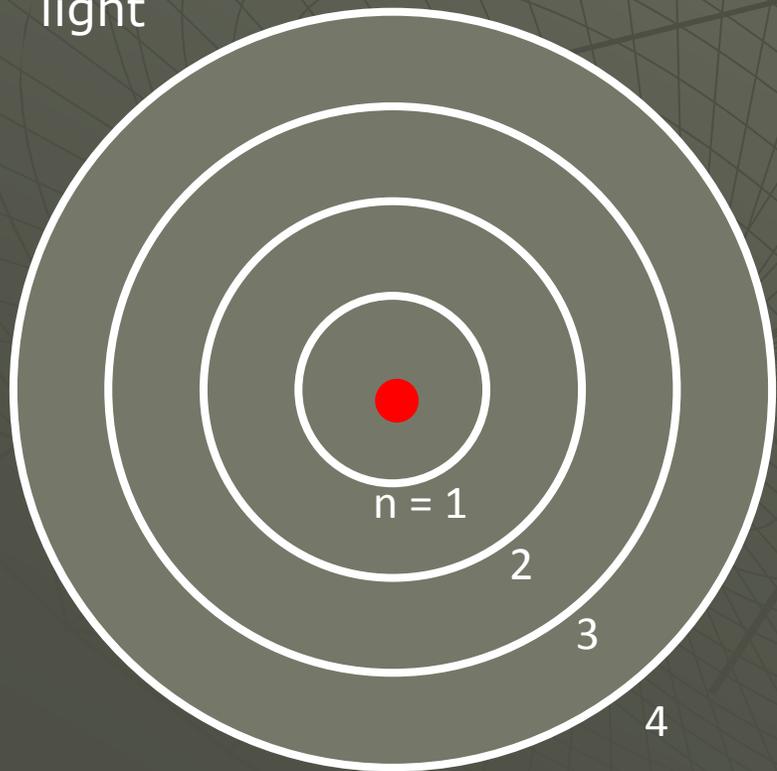
Absorption & Emission of Light

- (a) Photon strikes electron in ground state
- (b) Absorption: Photon *absorbed* as electron gains energy, moves to excited state
- (c) Emission: Electron falls from excited state to ground state, emits a photon of the same λ and energy



Orbit Energies

- Each orbit change requires an **EXACT** amount of energy
- These EXACT energies correspond to specific colors of light



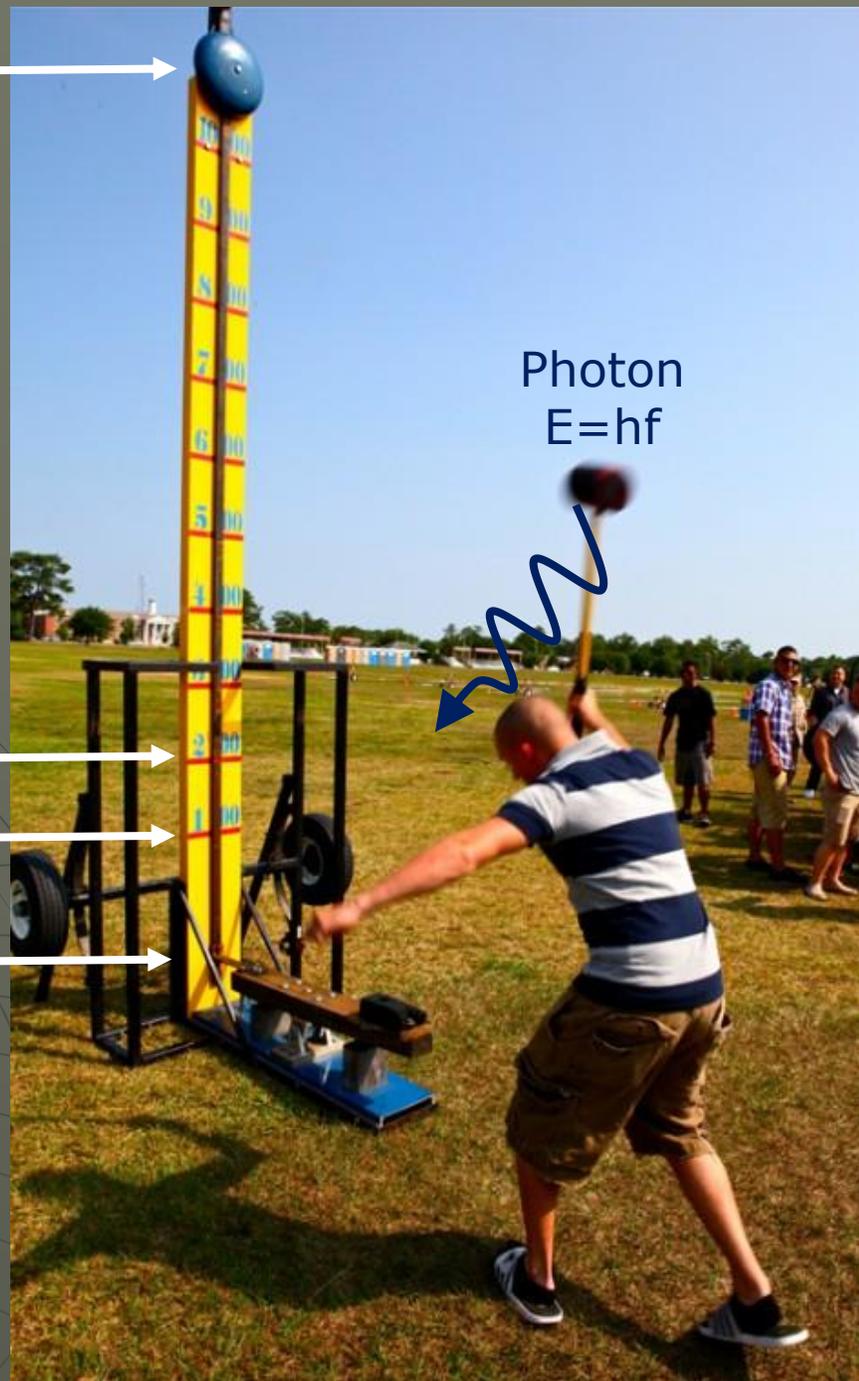
Higher energy ↑

2nd excited state $n=3$

1st excited state $n=2$

Ground State $n=1$

"Ionized" (electron escapes) $n=\infty$



Energies

- ◆ Gravity analog: a ball bearing held at a greater height has more energy ($PE = mgh$)
- ◆ Higher orbits (n is greater) = **higher energy**
- ◆ **Conservation of Energy**: same amount of energy absorbed must be released.
 - **Photon released has the same wavelength as the photon absorbed**

Energies

- ◆ Proton and electron *attracted* to one another with electrostatic force
 - Think magnets – really strong force
- ◆ Electrostatic force: $F \sim \frac{1}{r^2}$ (similar to gravity)
- ◆ Stable orbit: balancing forces
- ◆ What would happen to a stable orbit if the Mass of Earth increased.
 - A: move further away

Different Atoms

- H has 1 proton (plus charge)
 - Needs 1 electron (negative charge)
- He has 2 protons
 - Needs 2 electrons



$n=2 \rightarrow n=1$
H: 660 nm

He: 580 nm
 $n=2 \rightarrow n=1$



TPS

1. You see a yellow-orange streetlight in Laramie, which is a “sodium vapor” light. What are the electrons in the sodium doing as they emit the yellow light you see?
 - A. Moving toward the sodium nucleus
 - B. Moving away from the sodium nucleus
 - C. Moving to higher energy states
 - D. Electron magic

TPS

2. Which transition corresponds to the highest energy light being emitted?

A. 4 to 1

B. 2 to 1

C. 3 to 1

D. 4 to 3

3. Imagine the transition from $n = 4$ to $n = 1$ gives us violet light. Now let's say we witness the light emitted when an electron jumps down from $n = 6$ to $n = 1$. Hypothesize what kind of light this would be...

- A. Radio waves
- B. Microwaves
- C. Infrared
- D. Ultraviolet

4. A photon strikes an electron in the ground state. This photon's energy exceeds the maximum energy of any transitions (jumps) the electron can make. What happens?
- A. The photon is reflected
 - B. The electron stays in the ground state and just moves faster
 - C. The atom absorbs heat
 - D. The electron is ejected from the atom
 - E. "Hulk SMASH!"

LT

- Light & Atoms

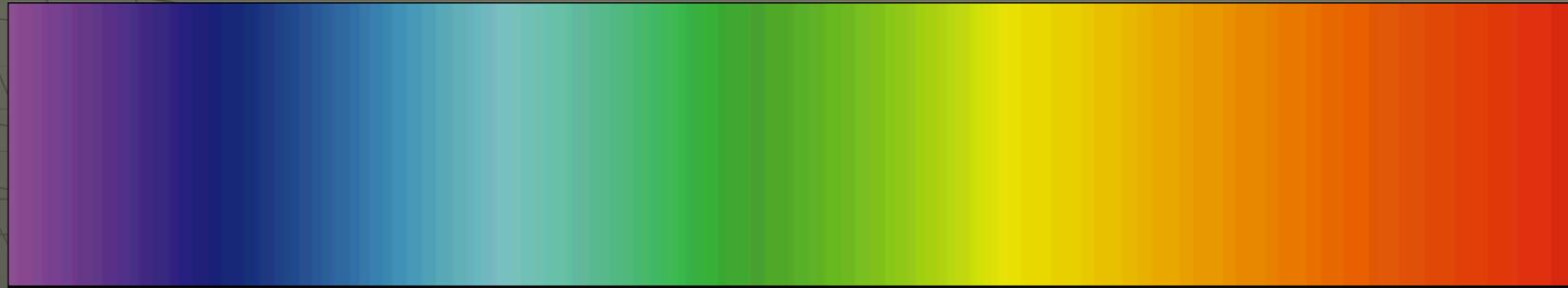
Spectroscopy

- ◆ Light we see from astronomical objects is filled with info that our eyes can't pick apart.
- ◆ **Spectroscopy:** The process of breaking light apart into its components, analyzing the colors present, their relative intensities, positions, etc.

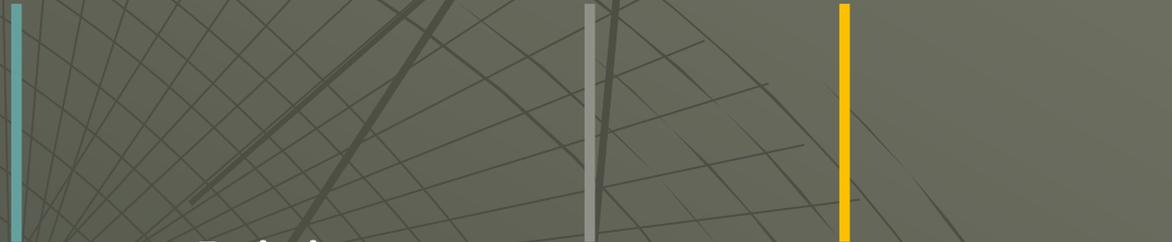
Kirchoff's Laws – Types of Spectra

- ◆ 3 types of spectra we observe from objects
- ◆ Conditions of material(s) generating the light determine which of the 3 types we see
- ◆ The 3 types are...

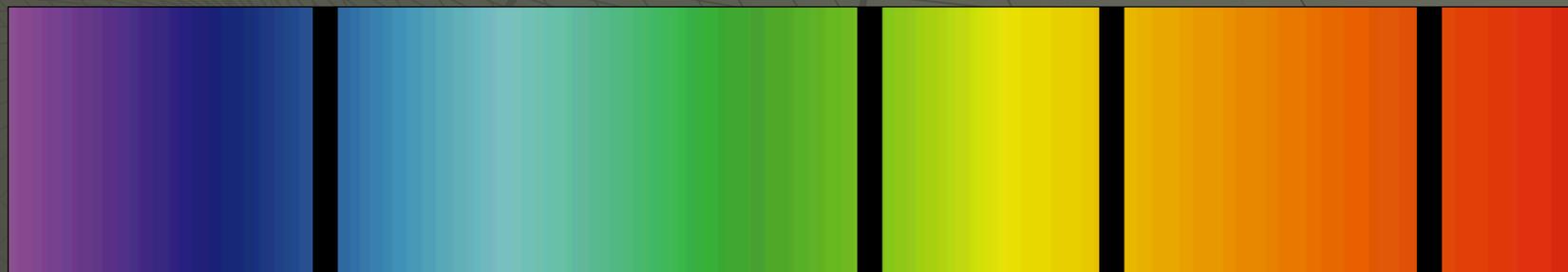
3 types of spectra observed



Continuous

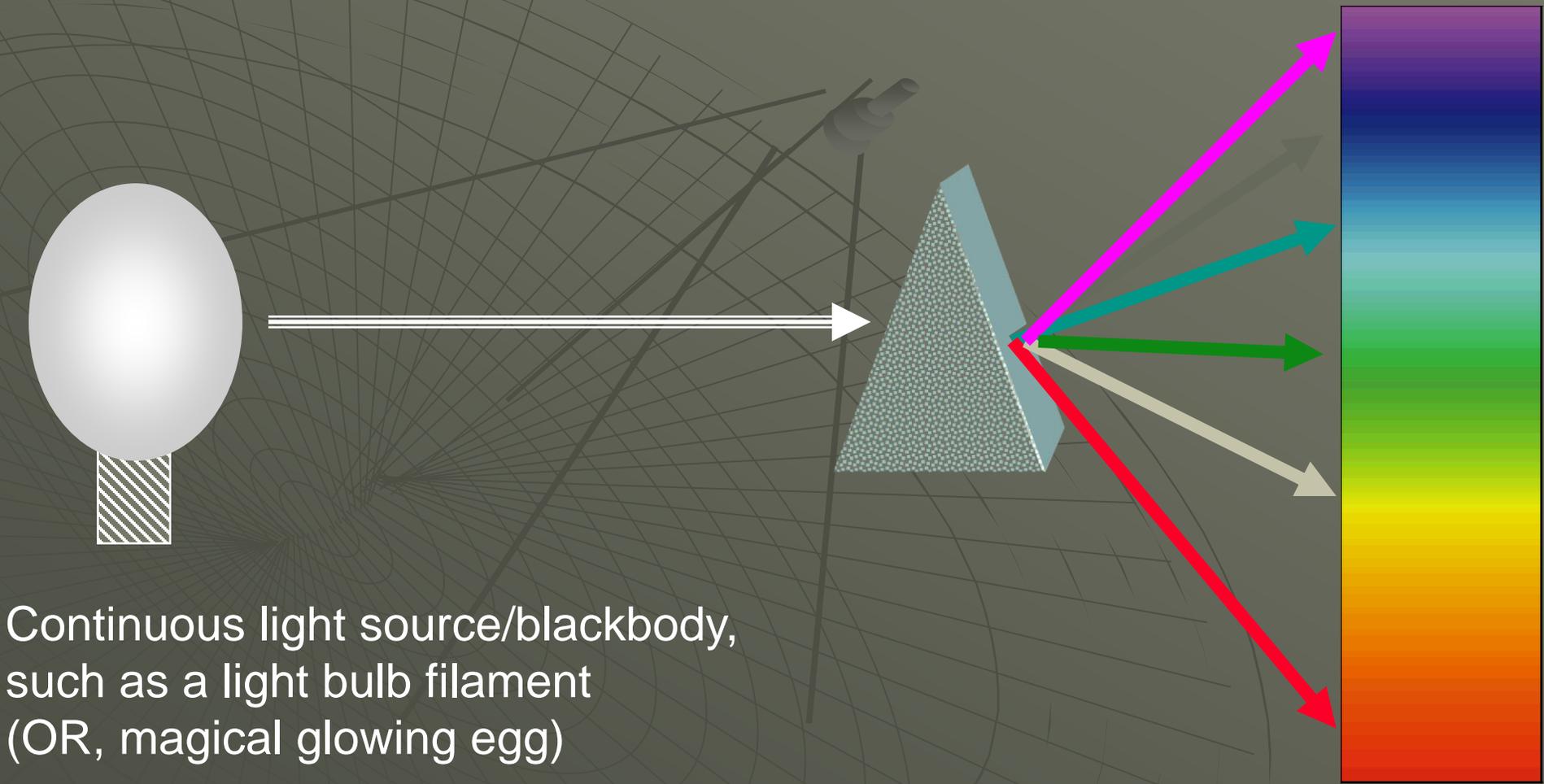


Emission



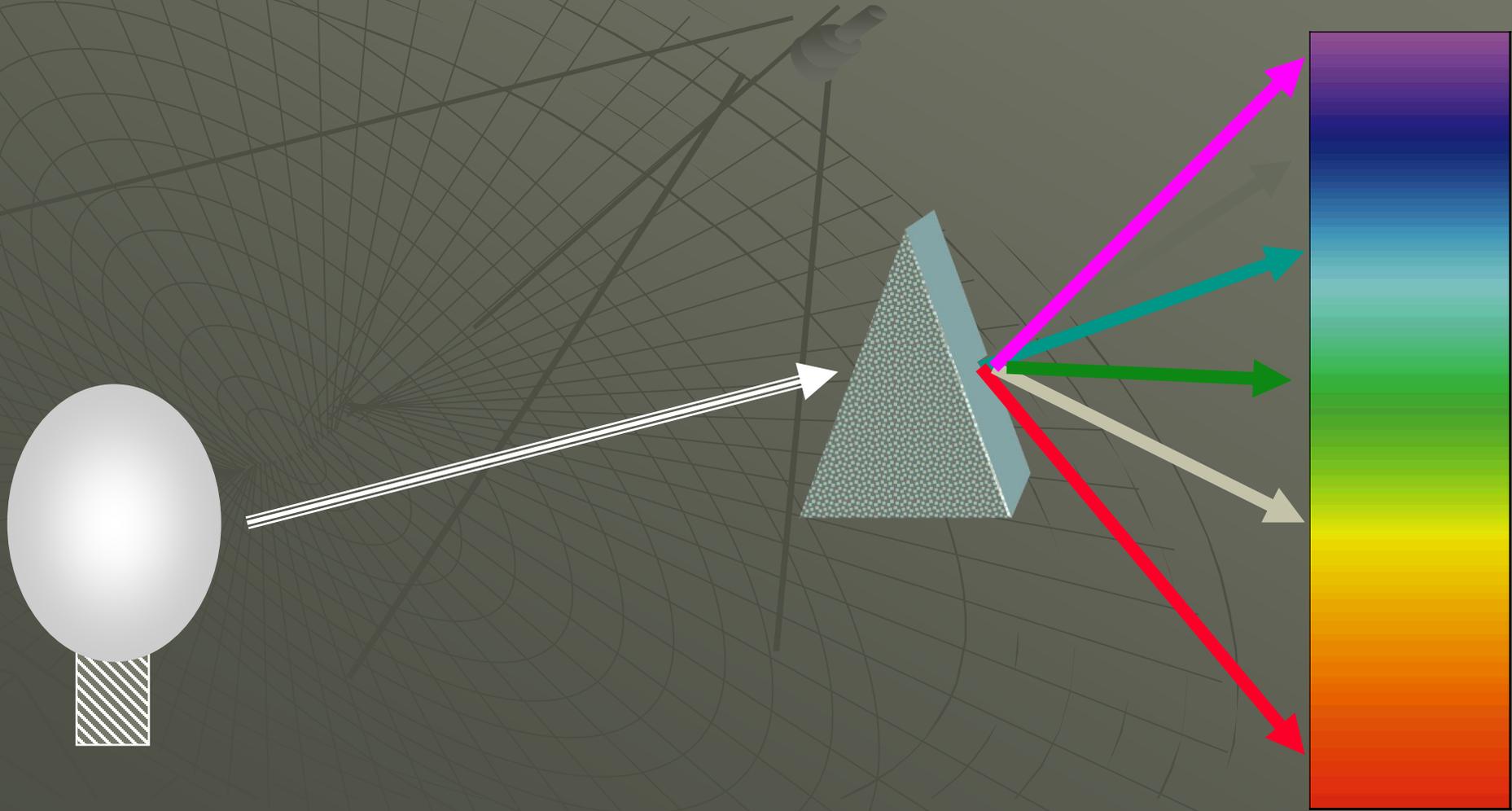
Absorption

Kirchhoff's Laws: # 1



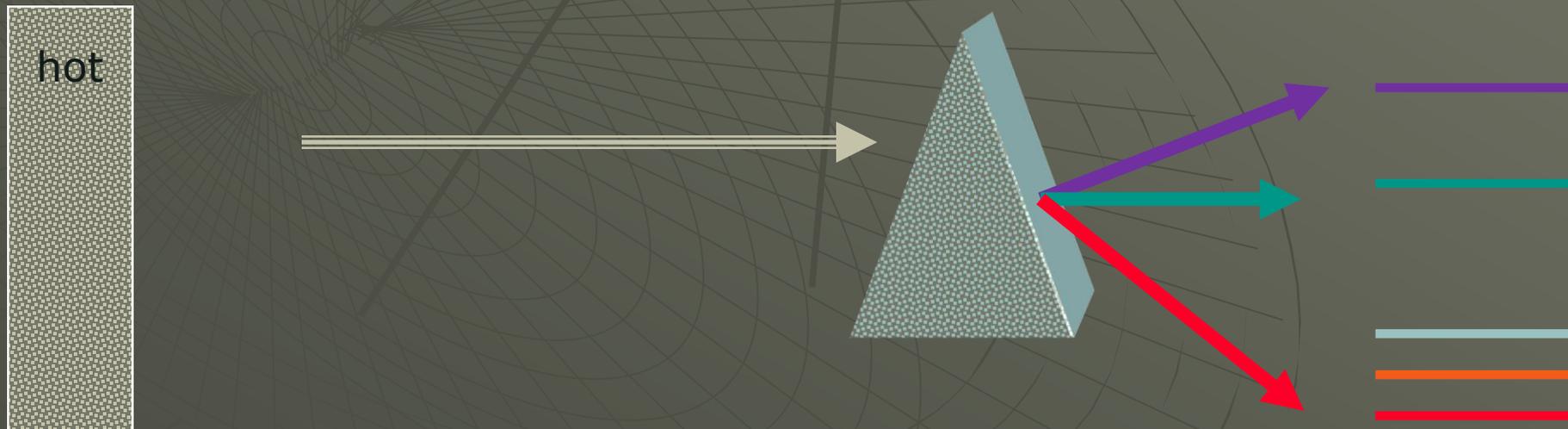
Continuous light source/blackbody,
such as a light bulb filament
(OR, magical glowing egg)

- ◆ A solid or a liquid, or a gas at a *high pressure*, emits a **continuous spectrum**. (a full “rainbow” of colors)

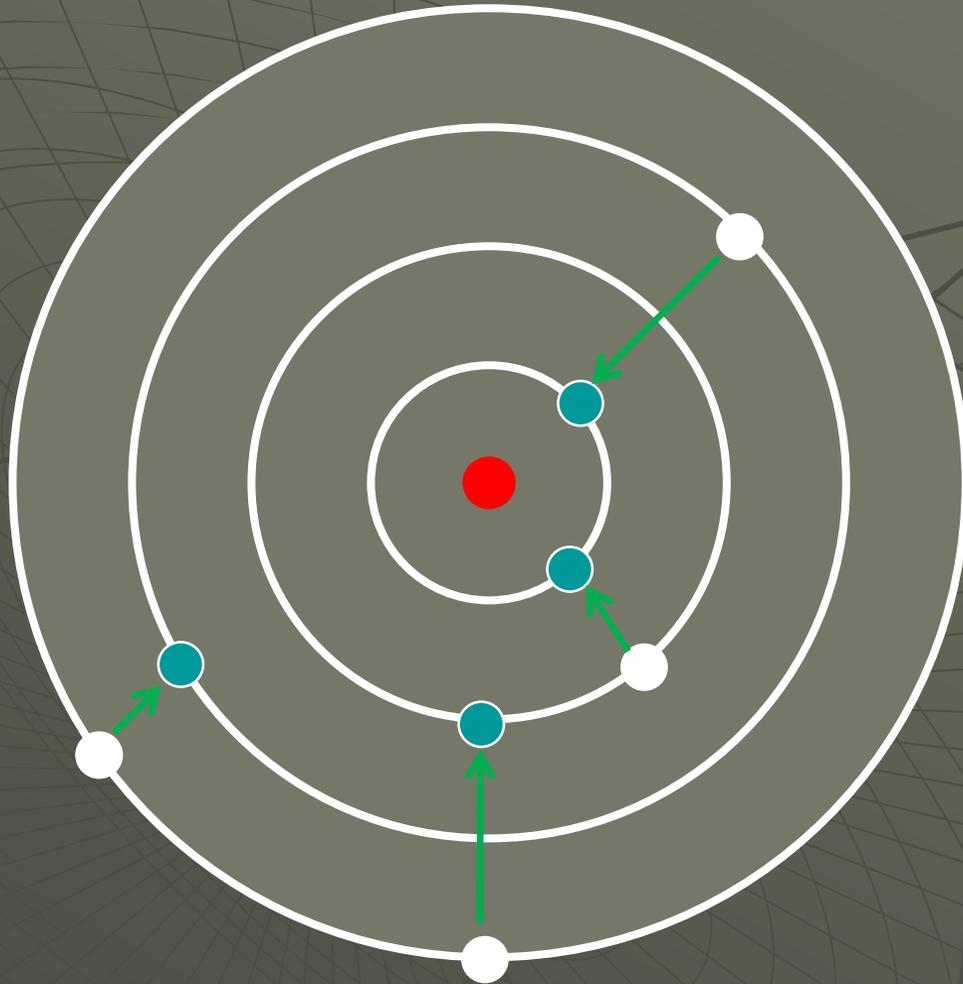


Kirchhoff's Laws: # 2

2. A hot, rarefied (low density) gas shows an **emission spectrum (very specific colors)**
 - ◆ Gas must be warm/hot – electrons in high energy states, and can emit light



Specific Energies...



Remember our Bohr Model. Only certain energy levels exist for the electron(s)

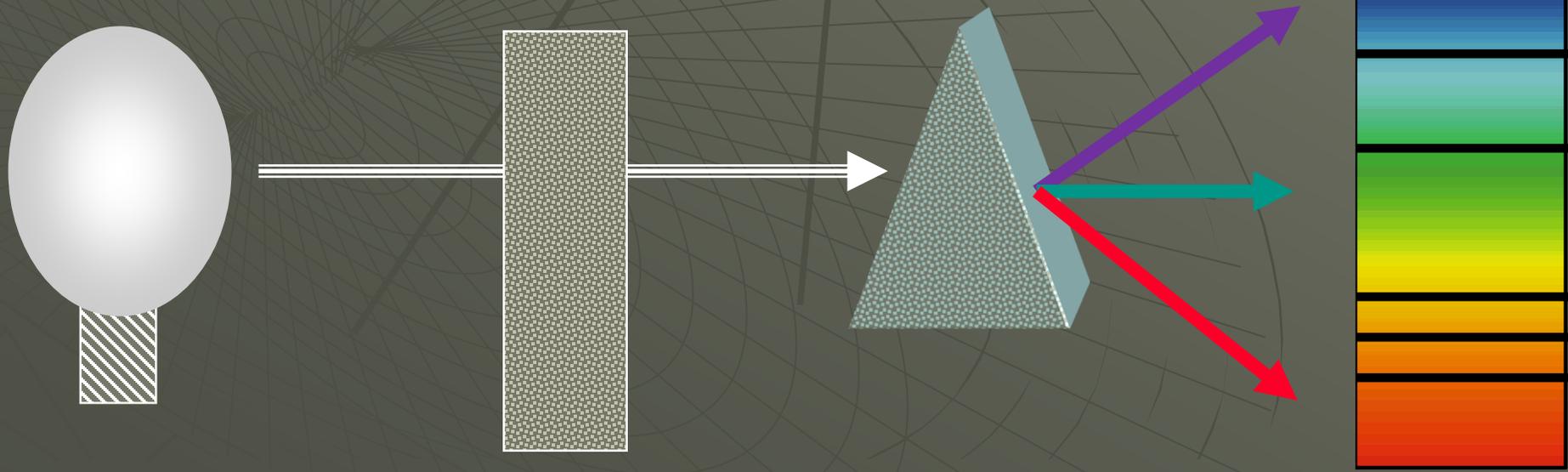
Each transition for an electron corresponds to a specific color of light

Now the cool part... (why spectra are a big deal for astronomers)

- ◆ These energy gaps are unique in each and every element in the universe.
- ◆ They are very specific, and can be measured with *extreme precision*.
- ◆ These spectra serve as chemical “fingerprint” for elements
- ◆ See poster in lab...

Kirchhoff's Laws: # 3

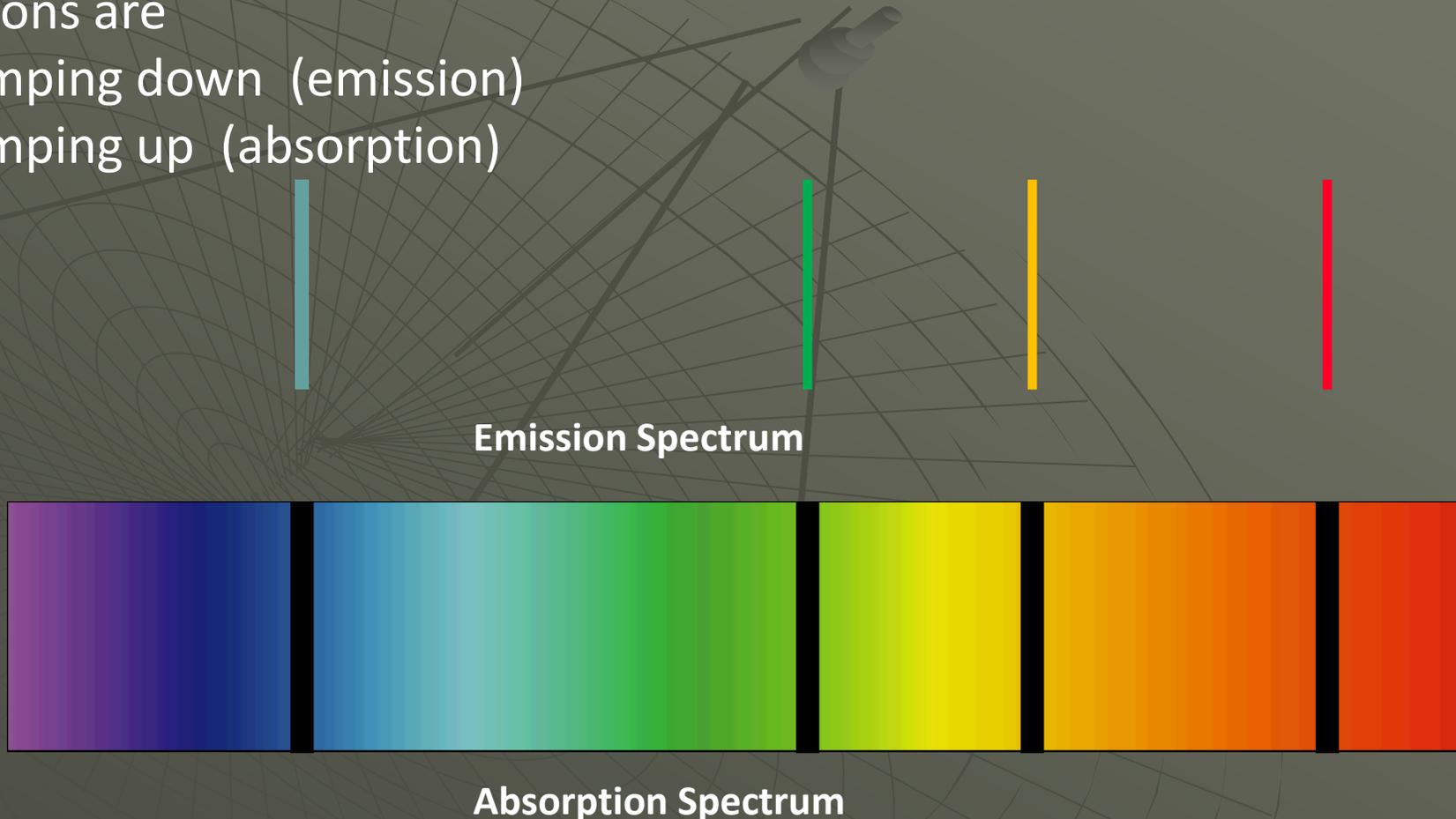
3. Light from continuous source, when passed through a cooler gas, may be absorbed at certain wavelengths, giving an **absorption spectrum**
- ◆ Gas must be relatively cool -- electrons in low energy states, and can absorb energy



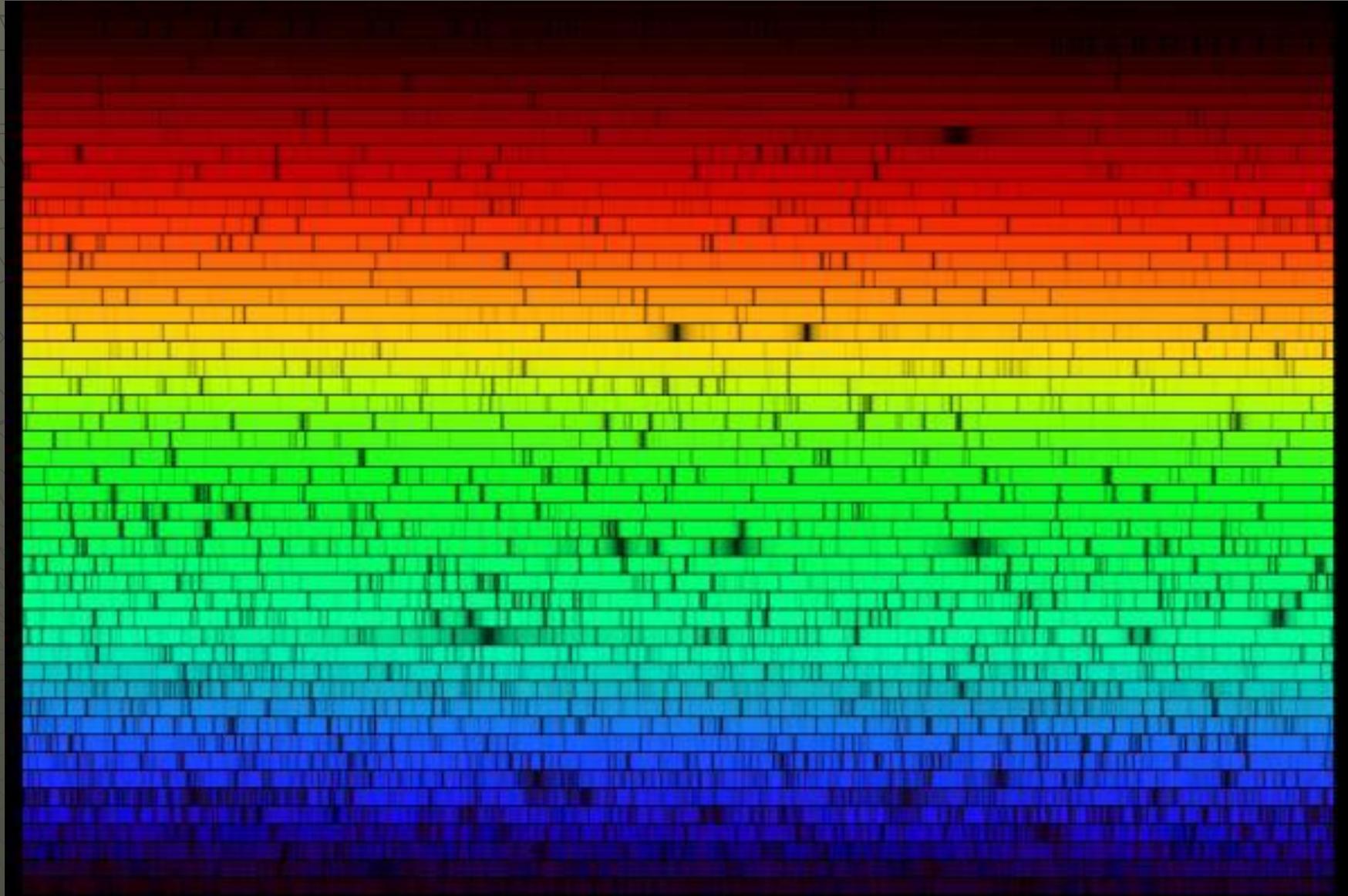
Not surprisingly...

A certain type of gas will *emit* light at very specific colors. It will *absorb* light at these exact same colors! Emission vs. absorption simply depends on if electrons are

1. jumping down (emission)
2. jumping up (absorption)



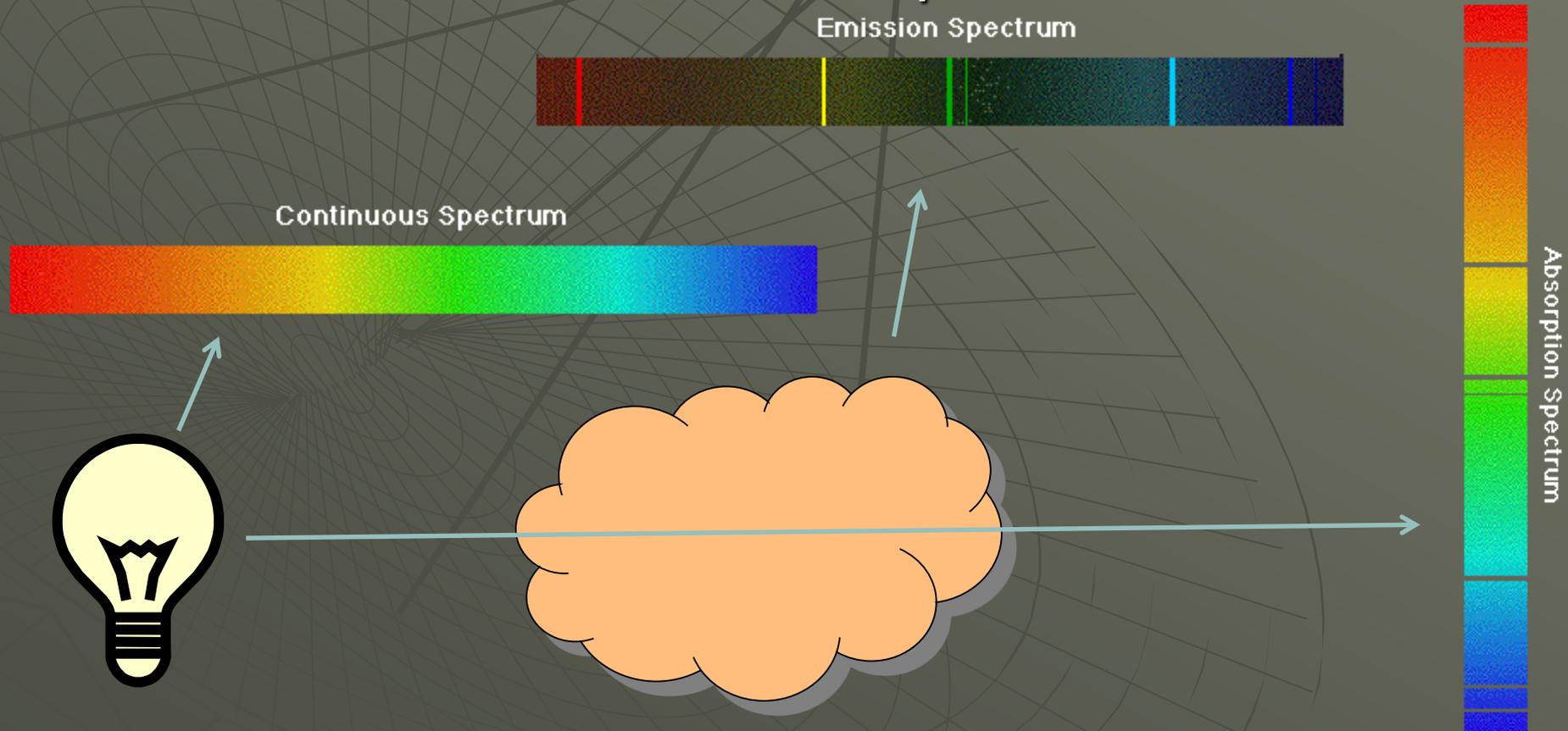
The Solar Spectrum (an absorption spectrum)



Using a spectrograph, we can resolve the sun's spectrum in great detail. Every single one of these dark lines is absorption from a specific element.

Kirchoff's Laws: Summary

- ◆ When do we see *emission lines*, and when do we see *absorption lines*?



LT

- ◆ Types of Spectra